



Hydrodynamics Experiments on the NIF: Scientific Accomplishments and Opportunities

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NIF/Jupiter User Group Meeting
Livermore, CA

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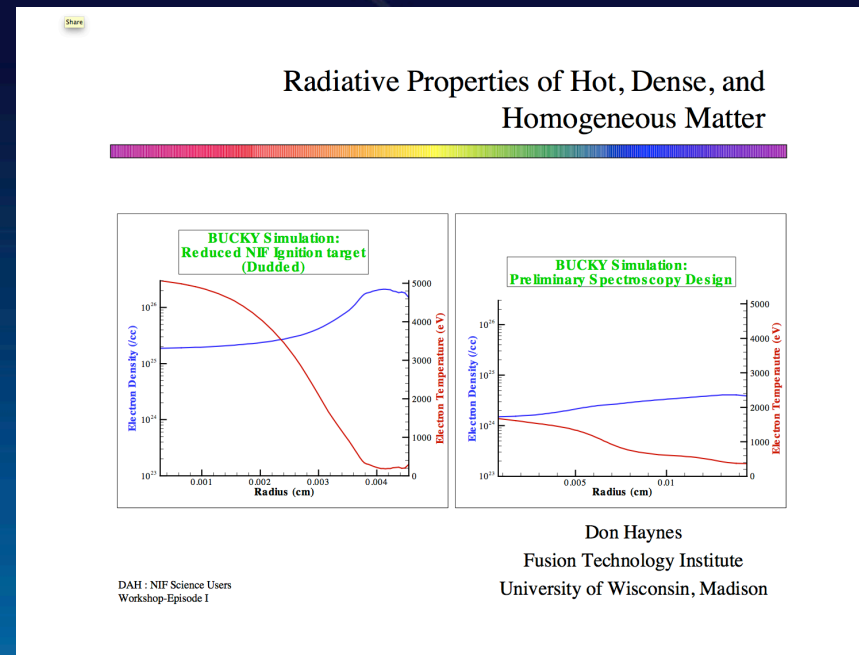
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It's been far too long.

- Back in 1999, at what may have been the first NIF Science Users Group Meeting, I discussed the difficulty of using the point design as a platform for line broadening studies (then the most important subject in the universe) and proposed an alternative design.
- Since then I've worked on IFE and on weapons.
- I am delighted to be back.



Presented at the National Ignition Facility Science Users Workshop, Pleasanton, CA, 5 October 1999

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Let's parse the title and derive an outline

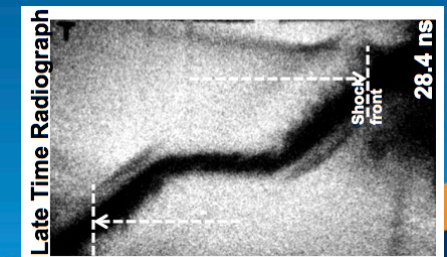
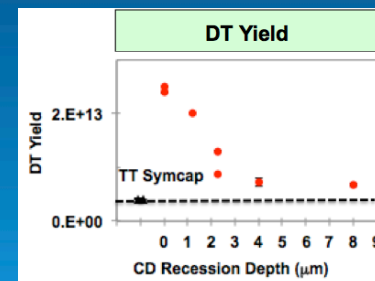
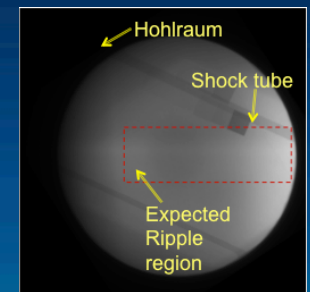
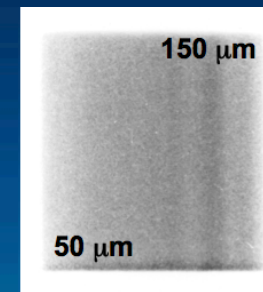
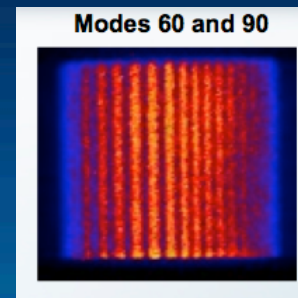
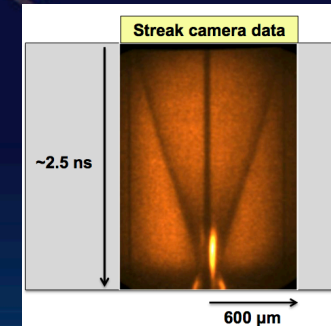
- Hydrodynamic Experiments on the NIF
 - Usefully limits the scope of what I need to cover in 30 minutes
- Accomplishments
 - The work of others (Bruce Remington, Alexis Casner, Vladimir Smalyuk, Carolyn Kuranz, Randy Kanzleiter, Forrest Doss, Ma, ConA team)
- Opportunities
 - One man's opinion

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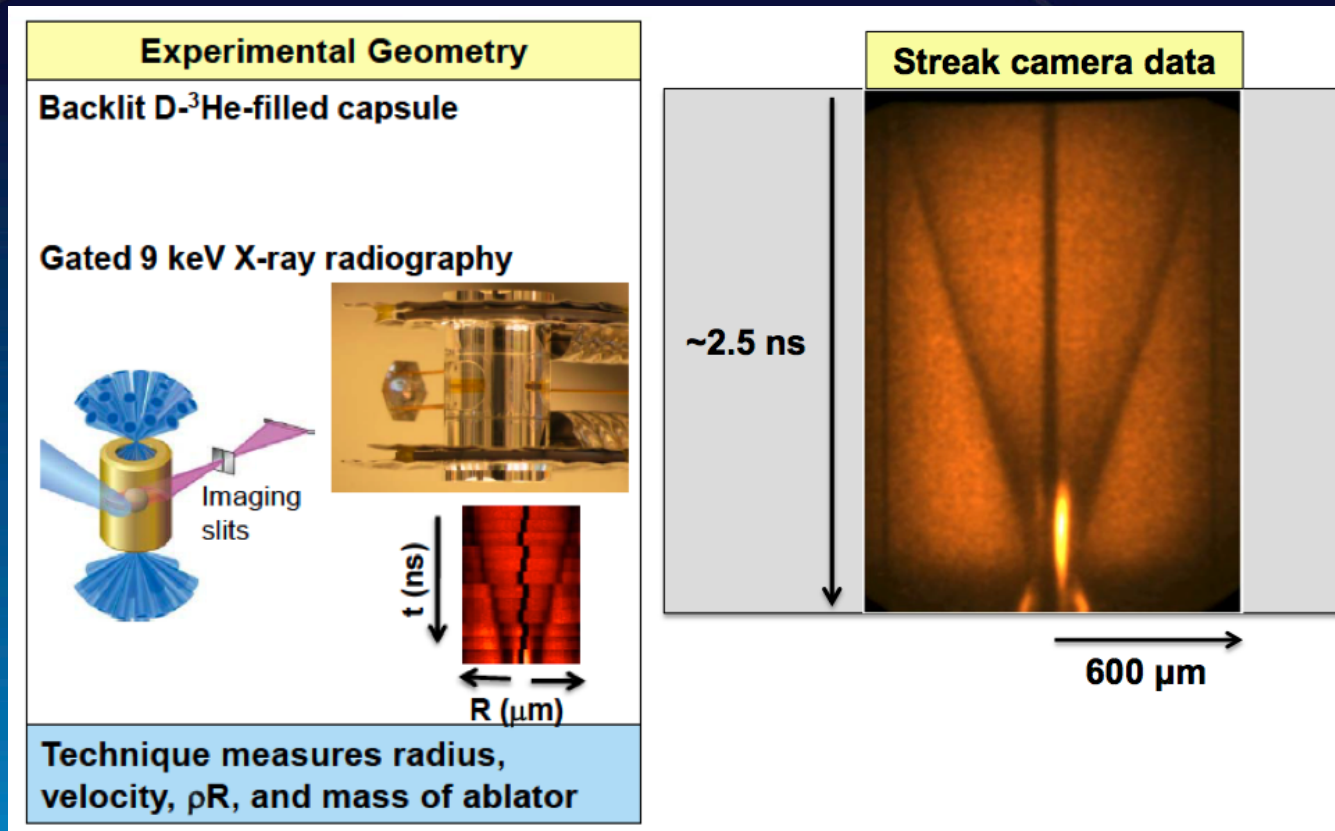
Hydrodynamic experiments at NIF

- Compressible Spherical Pistons
 - Convergent Ablators
- Instability growth factor studies
 - Hydro Growth Radiography Platform
 - ART, Rad SNRT
- Interfacial Mix and/or Turbulence
 - CD Mix
 - Shock/Shear



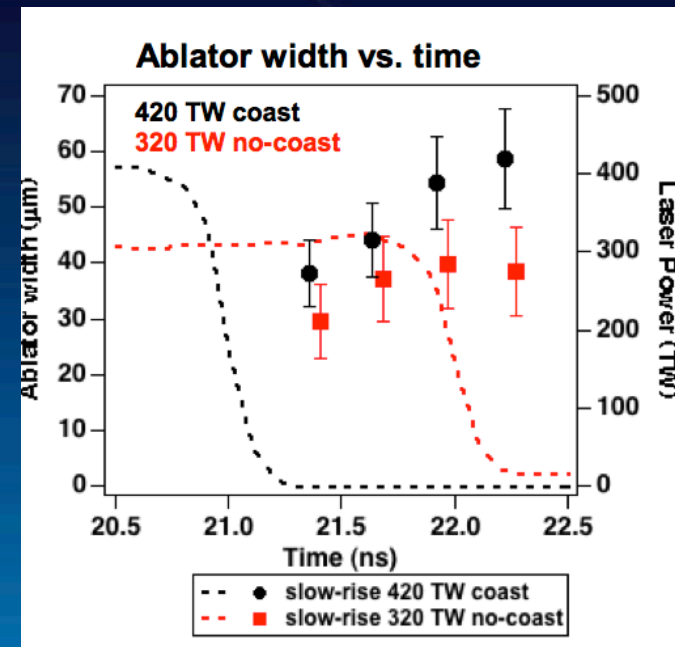
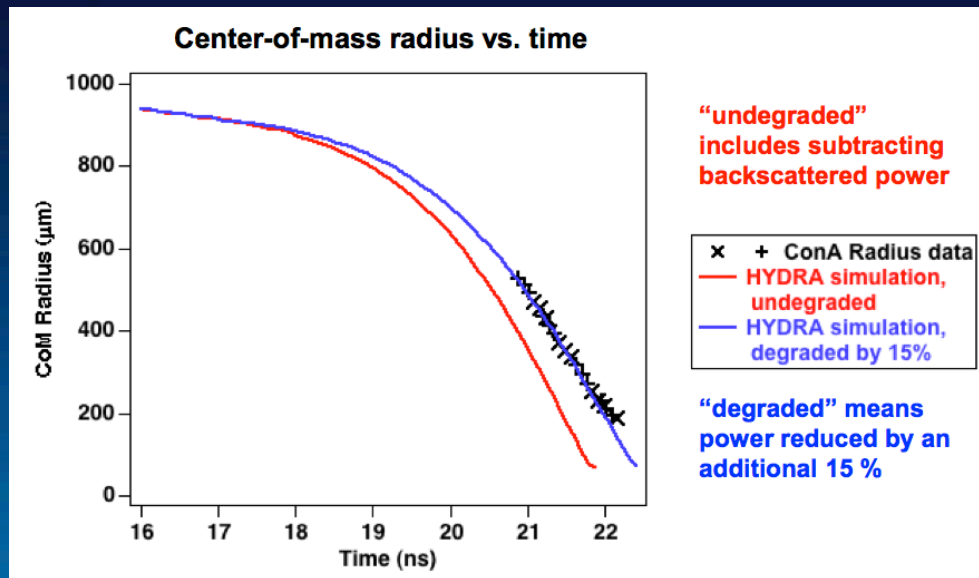
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Convergent ablator platform evolved to being an exquisite diagnostic of the behavior of the spherical piston on which all else relies in ICF



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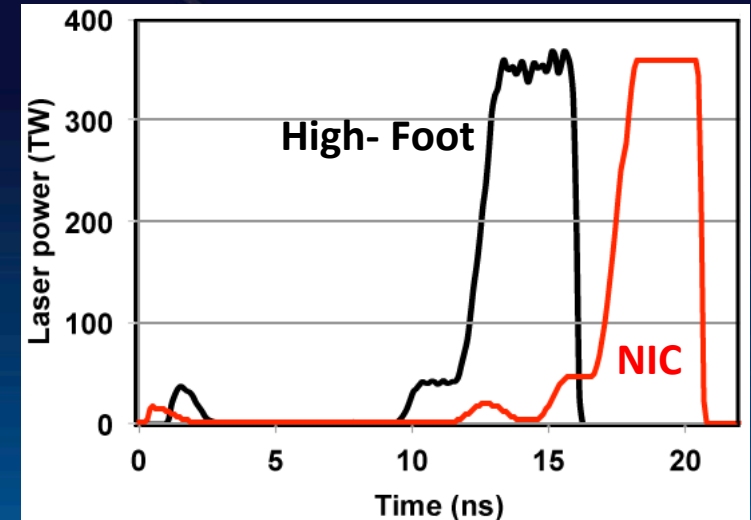
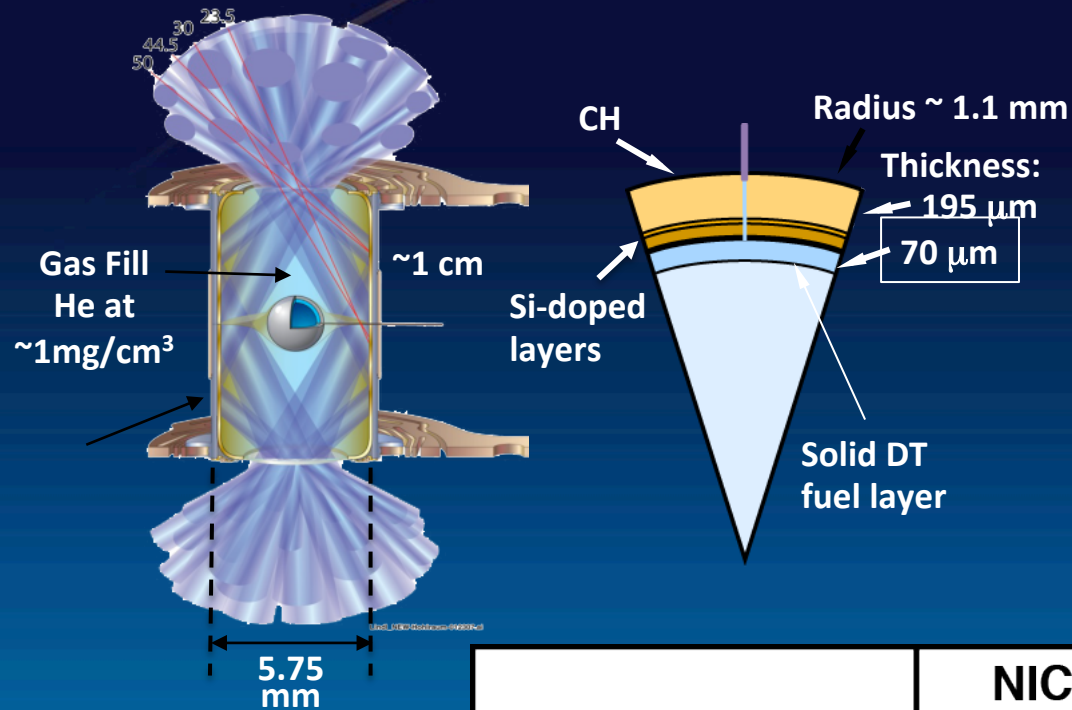
Convergent ablator platform evolved to being an exquisite diagnostic of the behavior of the spherical piston on which all else relies in ICF



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“High-foot” CH design uses same target with shorter laser pulse to reduce growth of surface perturbations



	NIC Low-foot		High-foot
Adiabat	~ 1.5	Increased to:	~ 2.5
In-flight aspect ratio, (IFAR)	~ 20	Reduced to:	~ 7
Convergence	~ 45	Reduced to:	~ 30

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If we are in a regime where amplitude = seed * GF then measuring the GF is important

- The high foot campaign trades off high gain potential for stability.
- The Hydro Growth Radiography platform allows us to directly test the underlying hypothesis: high foot => lower GFs.

$$kA^* \Delta u f - kv_a, \text{ early } RM$$

$$\sqrt{\frac{kAg}{1 + kL_\rho}} - kv_a, \text{ late } RT$$

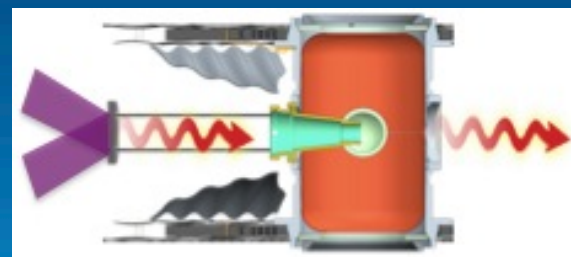
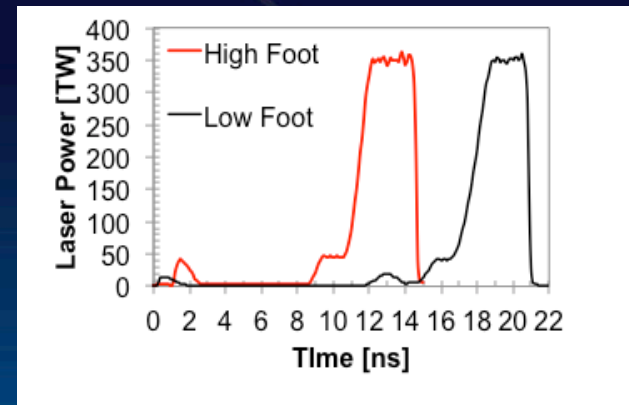
- Lower growth in Hi-Foot in RM phase because of shorter t_{RM} , higher v_{abl} , and fewer shocks
- Lower growth in Hi-Foot in RT phase because of larger $L = \text{grad}(\rho)/\rho$ and higher v_{abl}

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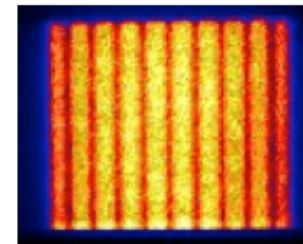


If we are in a regime where amplitude = seed * GF then measuring the GF is important

- Though the data requires some time shifting, by using amplitude as a function of wavelength current we can compare these apples and oranges.
- This platform is, in my opinion, a masterpiece of taking advantage of existing tech (slide cone, BL) to quickly answer an important question.

Radiographs with Low-Foot Drive

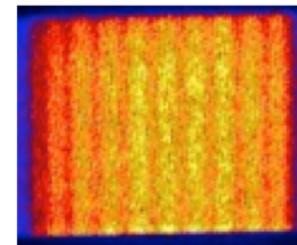
Mode 60



Shot N130602 @ 19.3 ns

Radiographs with High-Foot Drive

Mode 60



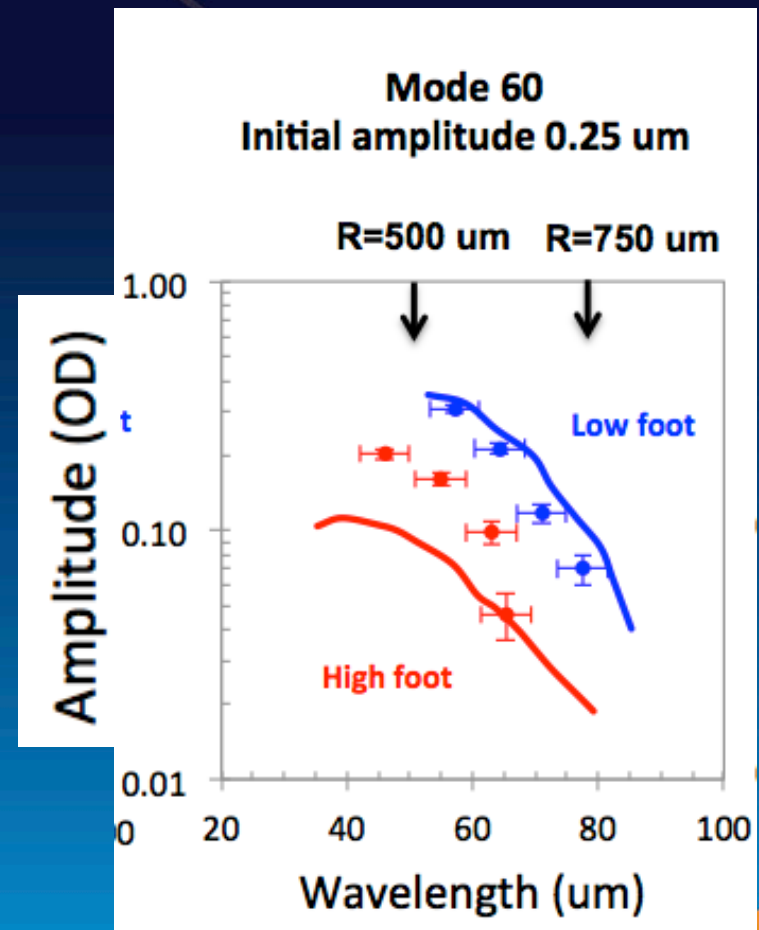
Shot N130702 @ 13.7 ns

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If we are in a regime where amplitude = seed * GF then measuring the GF is important

- Though the data requires some time shifting, by using amplitude as a function of wavelength we can compare these apples and oranges.
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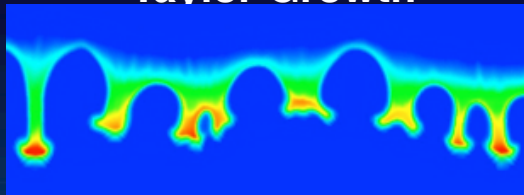


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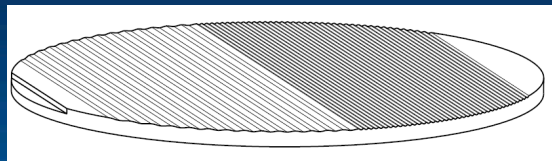
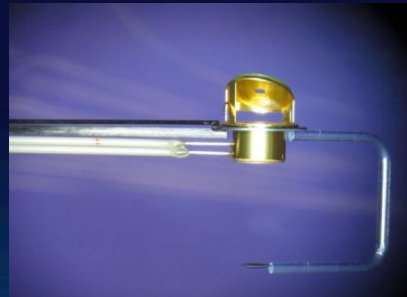


Ablative RT is to study the ablative RTi in the deep nonlinear regime and achieve bubble merger regime.

Highly-nonlinear Rayleigh-Taylor Growth



150 μm 50 μm ripples

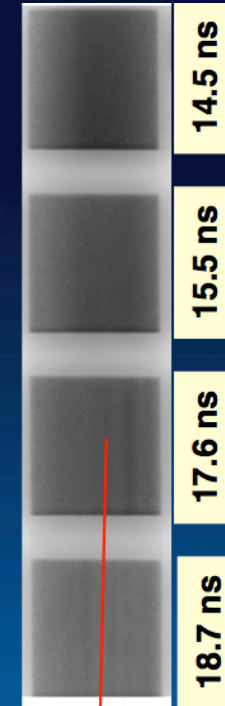


Driven from this side

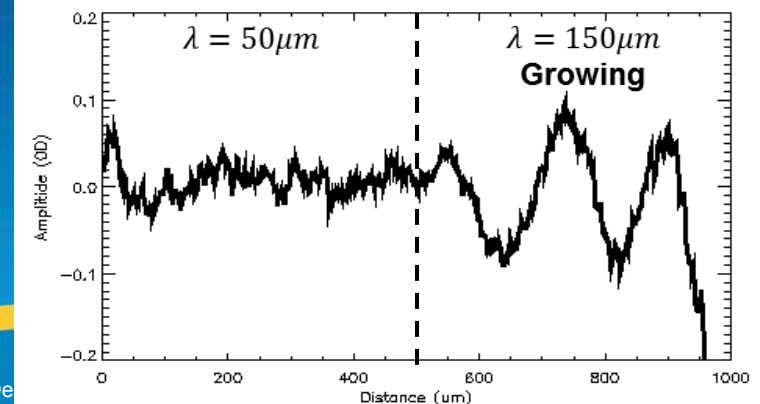
- Rarefaction wave transferred the modulations from the back surface to front surface.
- For the 150mm case the amplitude of the od modulations was significantly greater than postshot simulations.

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N131008 Primary rGXD1 data



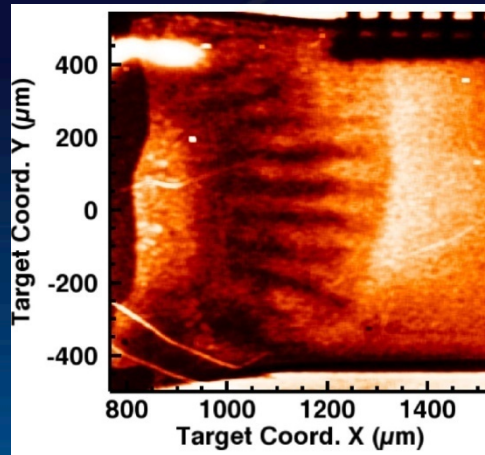
Lineout through slit





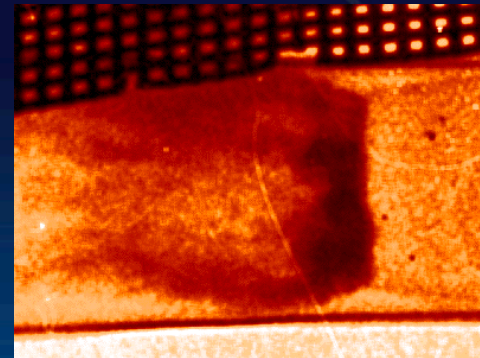
NIF will allow the exploration of the combined effects of radiative shocks and instability growth.

Supernova relevant hydrodynamics

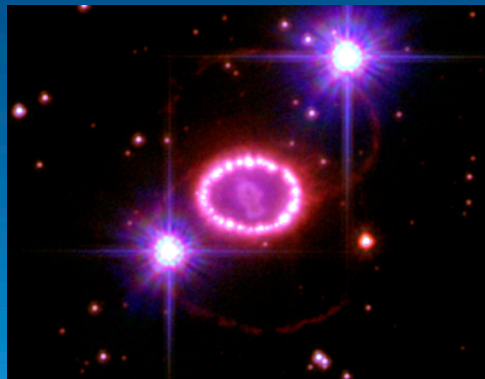


Scaled model of instabilities at H/He interface of SN1987A

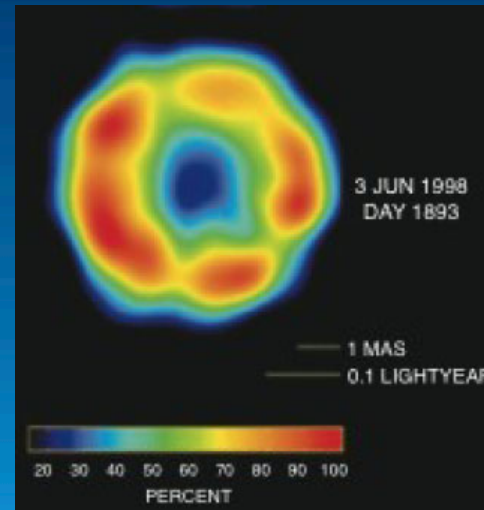
Supernova relevant radiative shocks



20x shock compression by radiative losses



SN1987A, a core-collapse, blue supergiant supernova (HST)



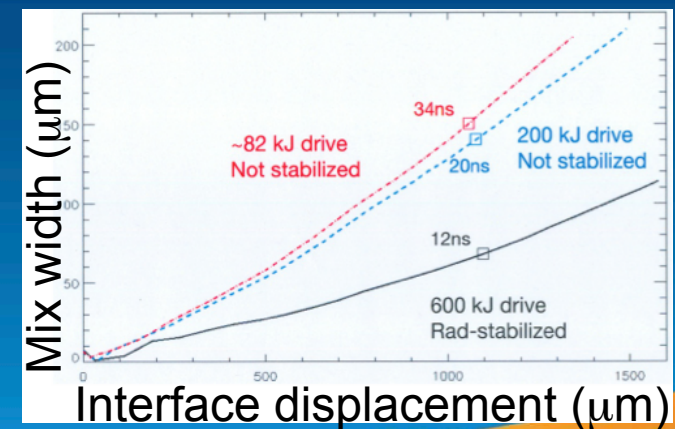
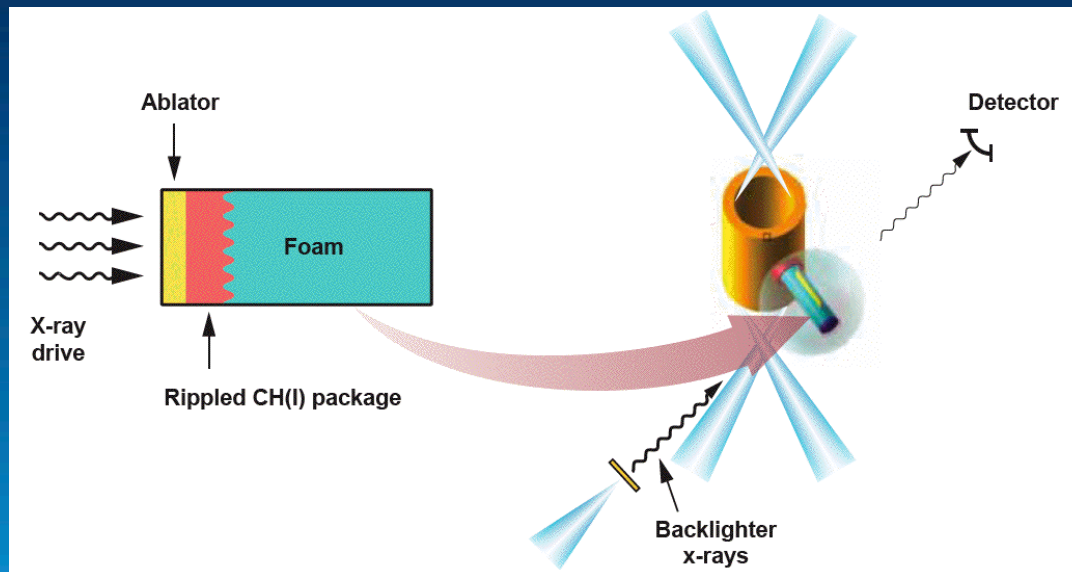
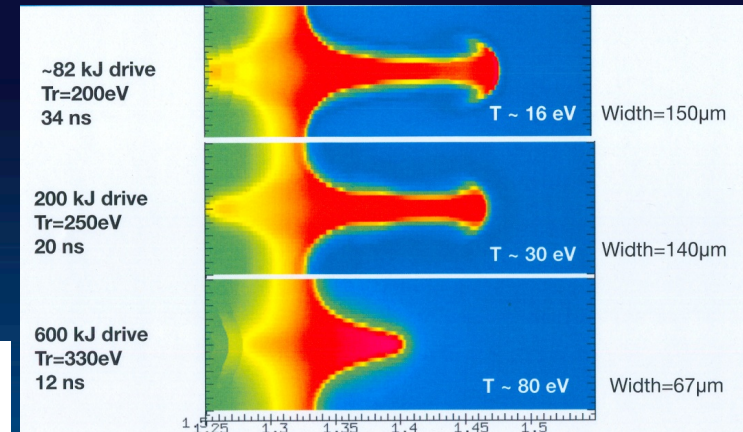
SN1993J, structure may be due to radiative collapse (Bartel, Science, 2000)

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As radiation becomes important, the qualitative nature of shock waves and instability growth also changes.

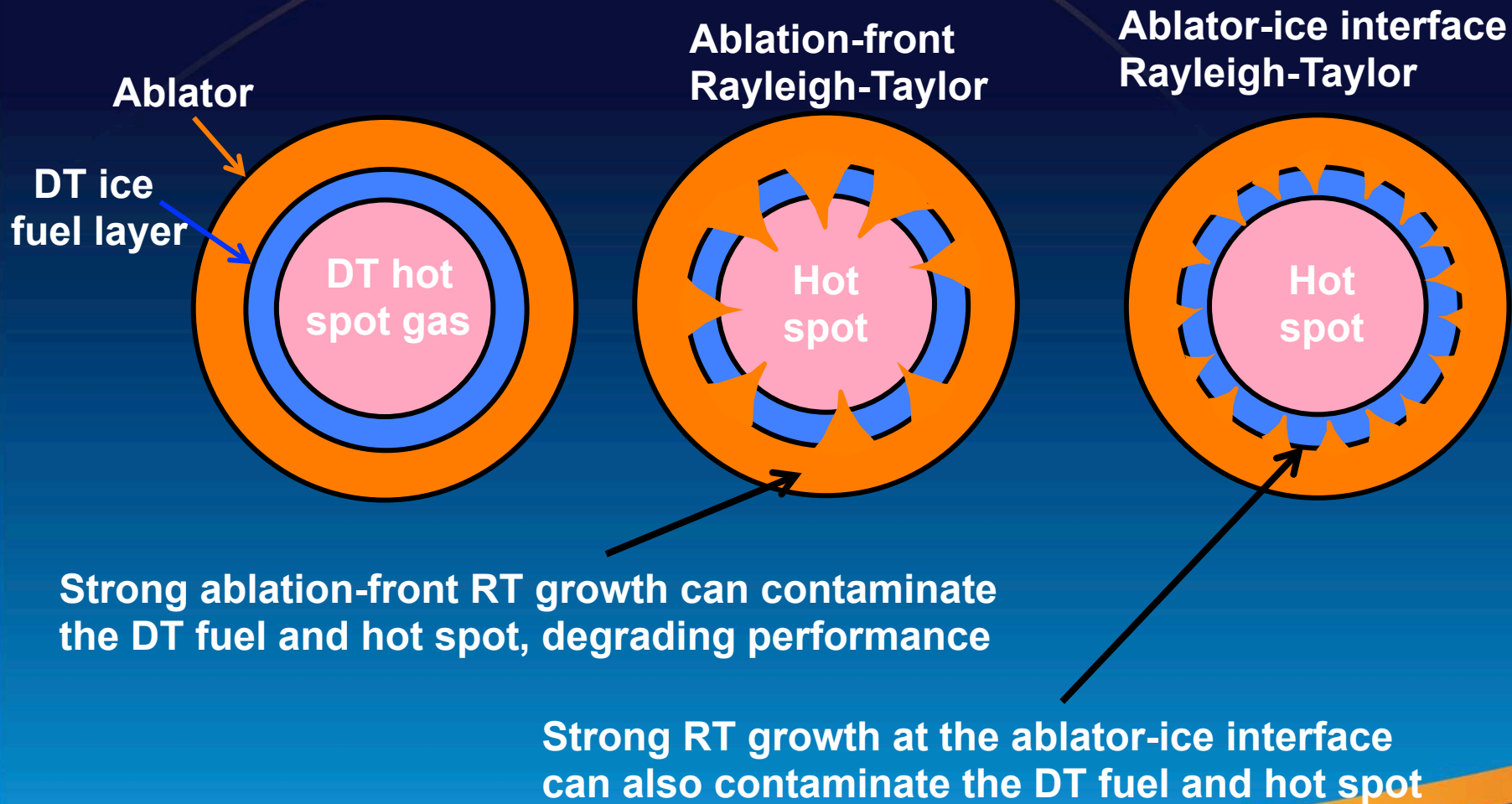
- Initial NIF experiments demonstrated a need to reduce background significantly. This work is underway.



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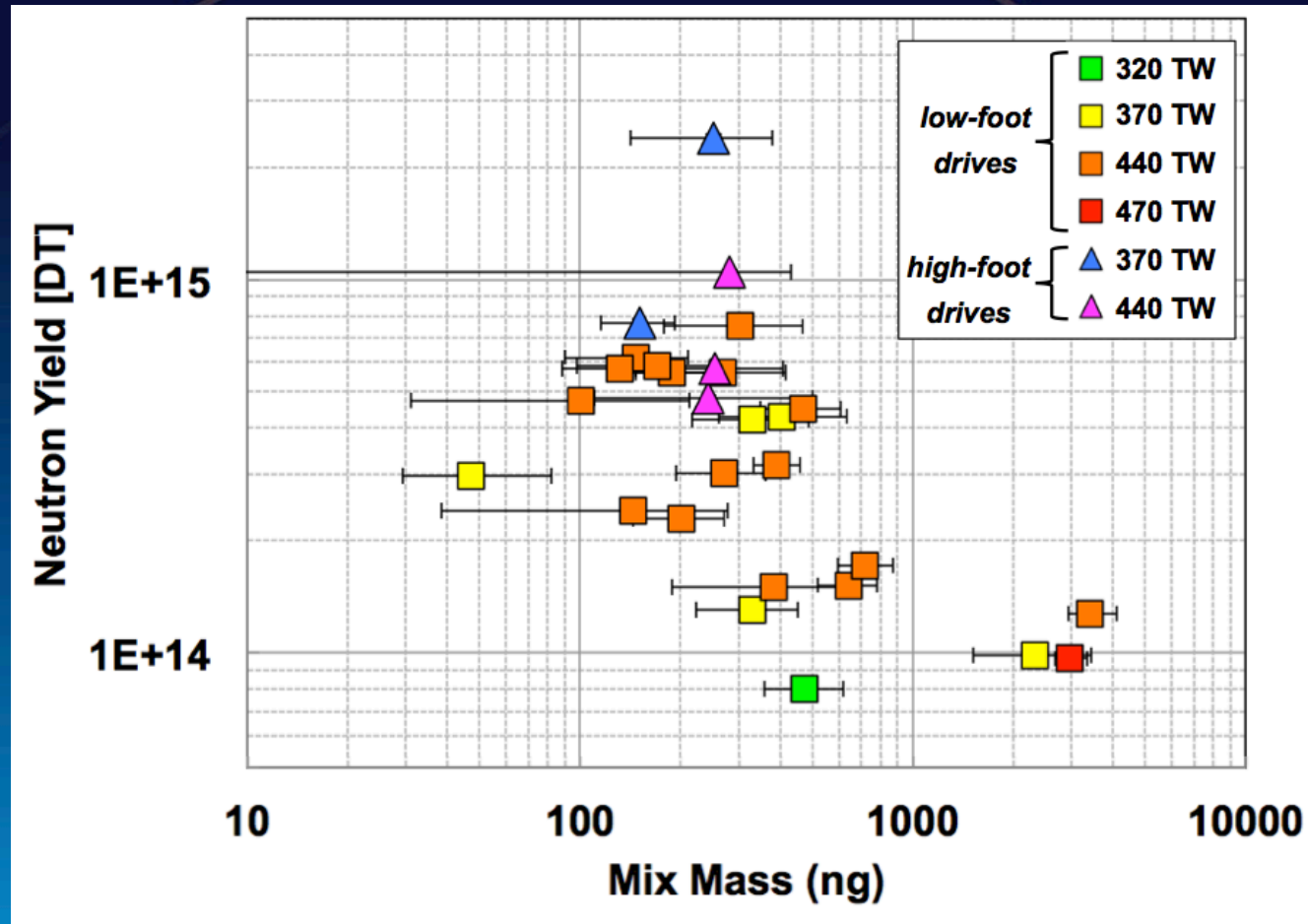
Bruce Remington is leading an international team investigating the physics of mix and ways to mitigate its deleterious effects on ignition.



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Before collecting experimental data, the low foot point design was thought to be robust to ~90ng of mix. If the xray/neutron yield ratio technique is as accurate as it is precise, nature disagreed.

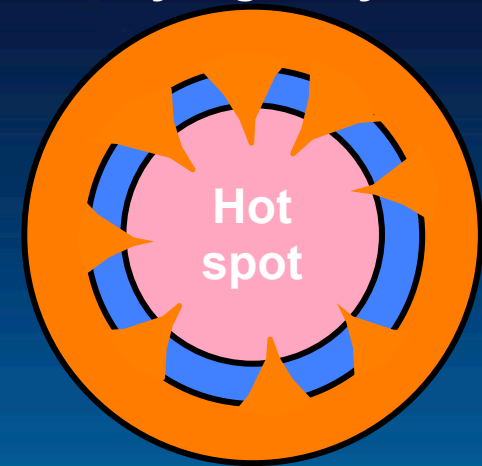


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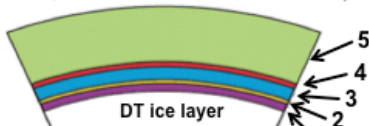


Strong ablation-front RT growth clearly contributes to mix, unless there is another mechanism to make the initially-distant-from-the-core Ge light up. Interestingly, there is no significant Cu detection from the innermost ablator layer.

**Ablation-front
Rayleigh-Taylor**

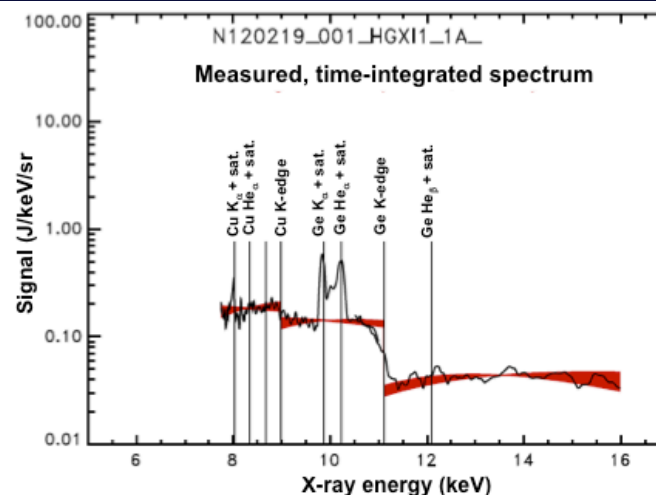


Tri-doped CH ablator (N120219)



Layer	Dopant (atm. %)
1	Cu(0.14%), Si(0%)
2	Ge(0.20%), Si(0.87%)
3	Ge(0.20%), Si(1.64%)
4	Si(1.15%)
5	None

S. P. Regan et al., "Hot-spot mix in ignition-scale ICF targets" PRL, 111, 045001 (2013).



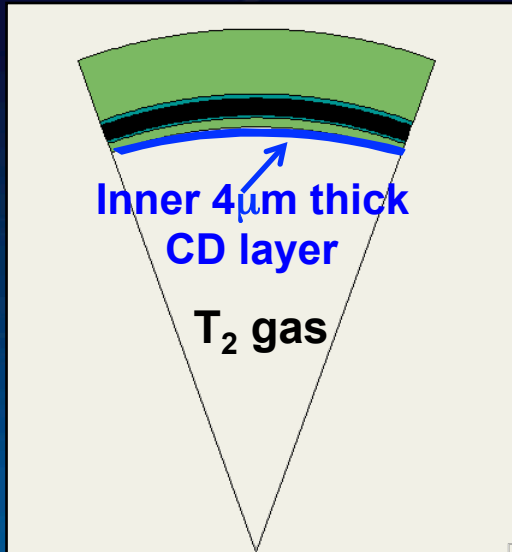
Note that the Cu (in the ablator layer closest to the fuel) didn't light up. Weird.

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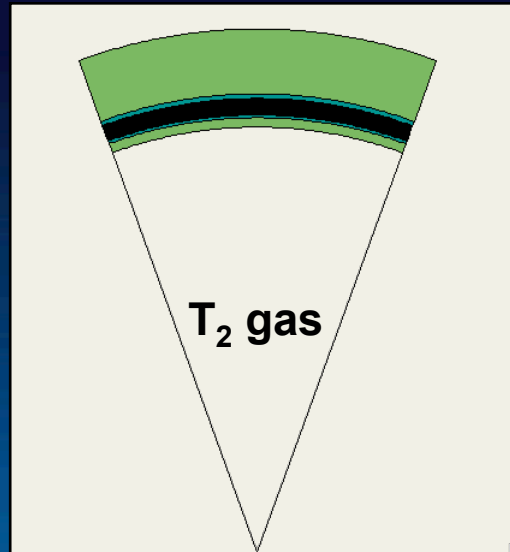
Nil desperandum, fans of ablator/fuel interfacial mix. Warm CD shells filled with (mostly T) show above 'background' DT, indicating for this system the first several microns of the ablator/ice do mix with the fuel 'atomically'.

CH shell with CD layer



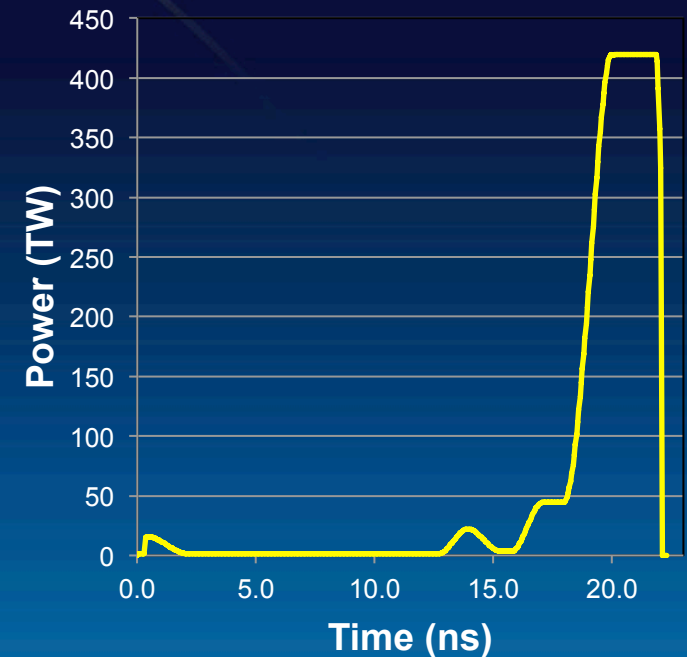
Mix measurement
 $D + T \rightarrow He + n$

CH shell



D contamination
measurement

Laser Drive 1.5MJ, 435TW



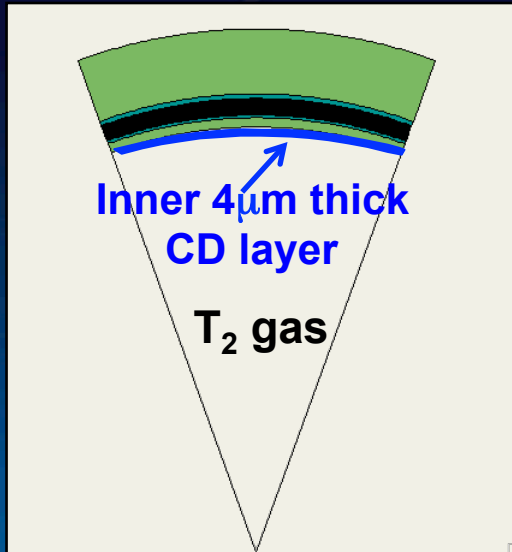
- CD-Mix experiments measure DT yield due to ablator-gas atomic mix
- Shells filled with 11.05 mg/cc T_2 gas at 32K including 0.1% D contamination

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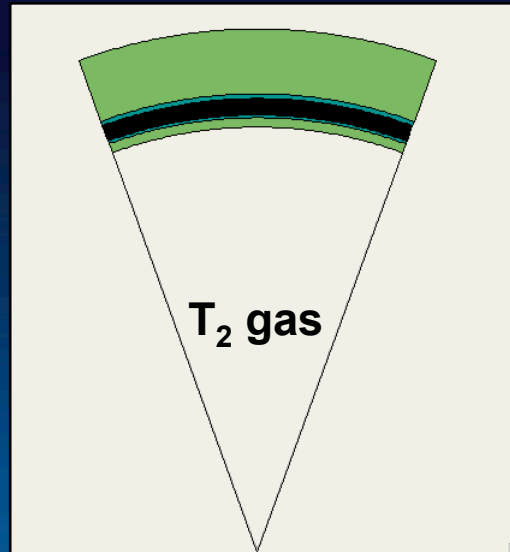
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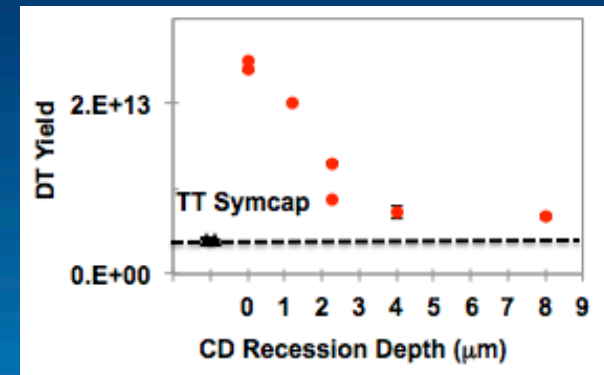
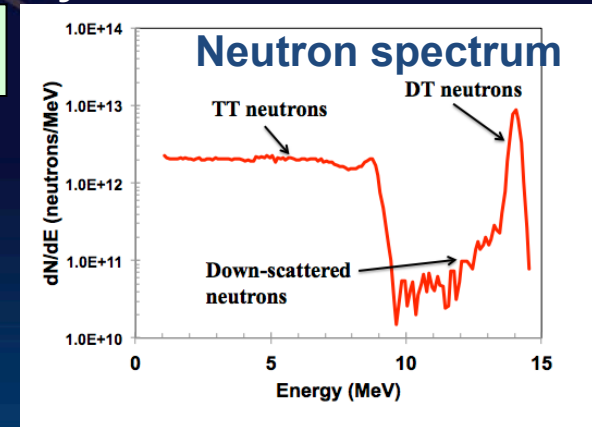
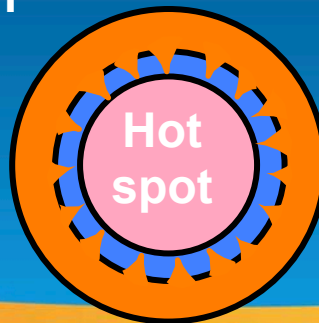


Mix measurement
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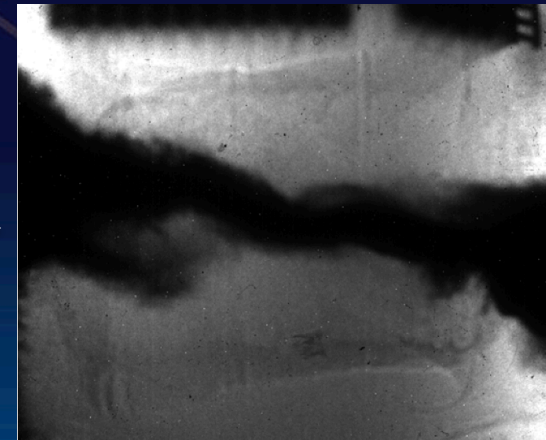
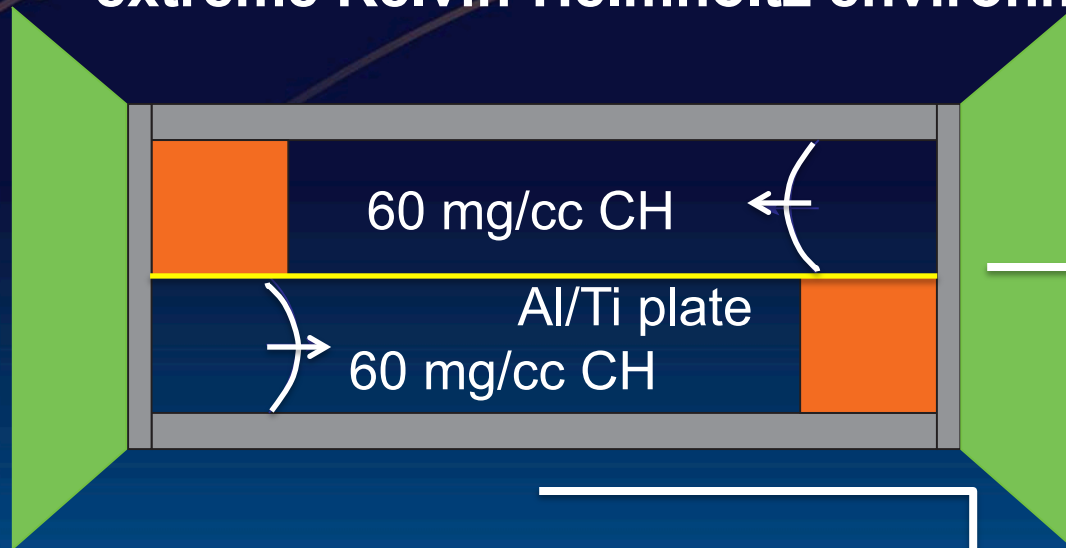


- CR ~ 20
- Surrogacy issues?

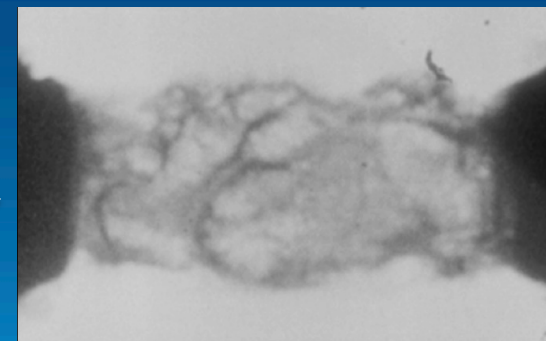
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Counterpropagating shear experiment creates extreme Kelvin-Helmholtz environments



- Top right: edge-on view measures tracer mix width.
- Bottom right: transverse view is used to image developing turbulence in the tracer plane.



F. W. Doss et al, "Instability, mixing, and transition to turbulence in a laser-driven counterflowing shear experiment," *Phys Plasmas*, **20**, 012707 (2013)

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F. Doss, L. Welser-Sherrill, K. A Flippo, E. Loomis, J. Fincke, 2013

Operated by Los Alamos National Security, LLC for the U.S. Department of Energy's NNSA

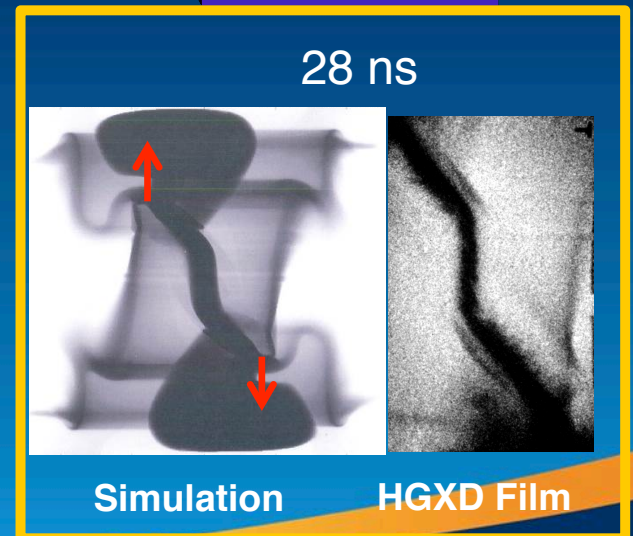
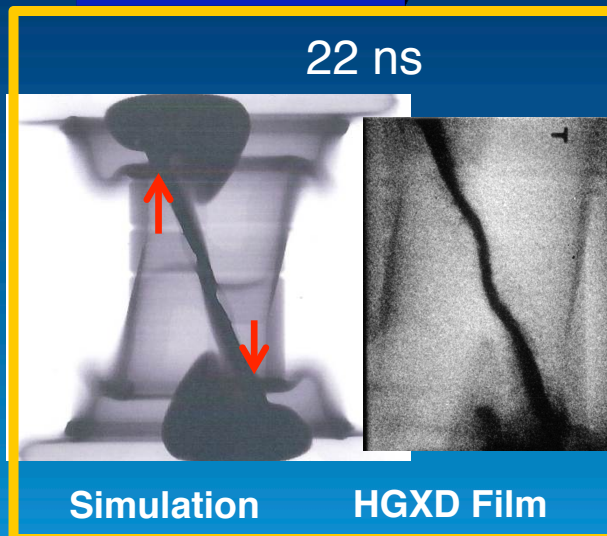
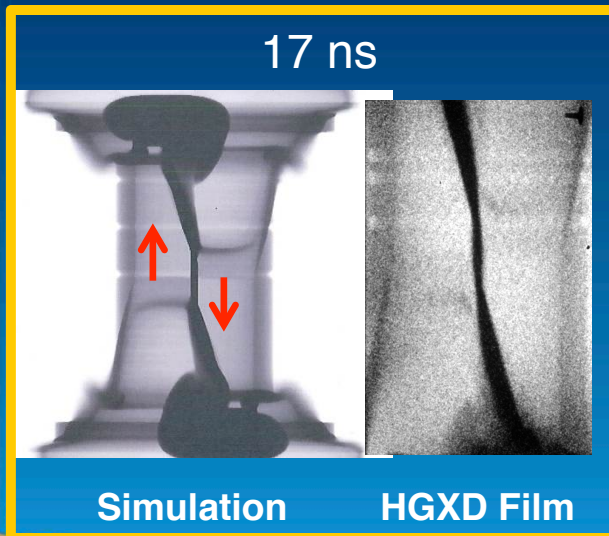
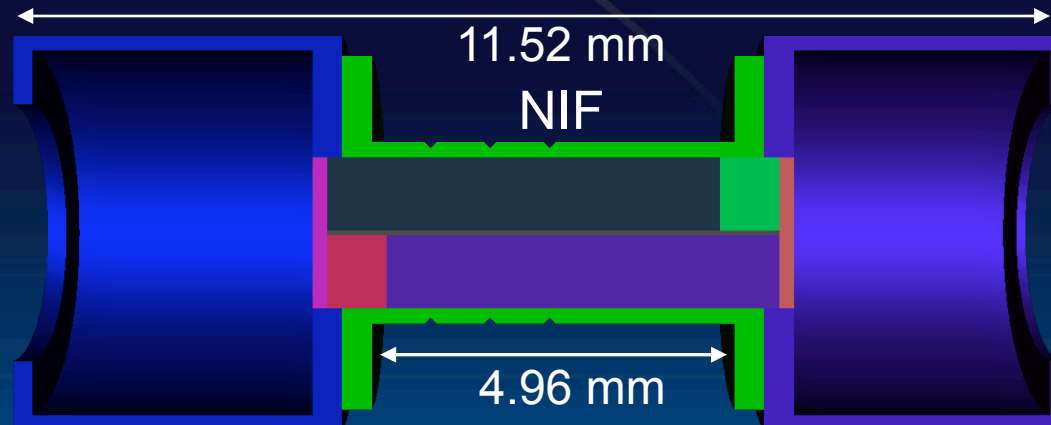
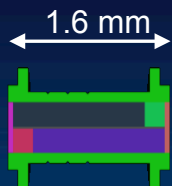


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The NIF platform allows connection to the regime explored at Omega.

Omega

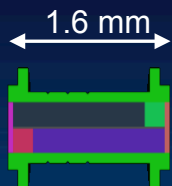


↓ shocks ↑

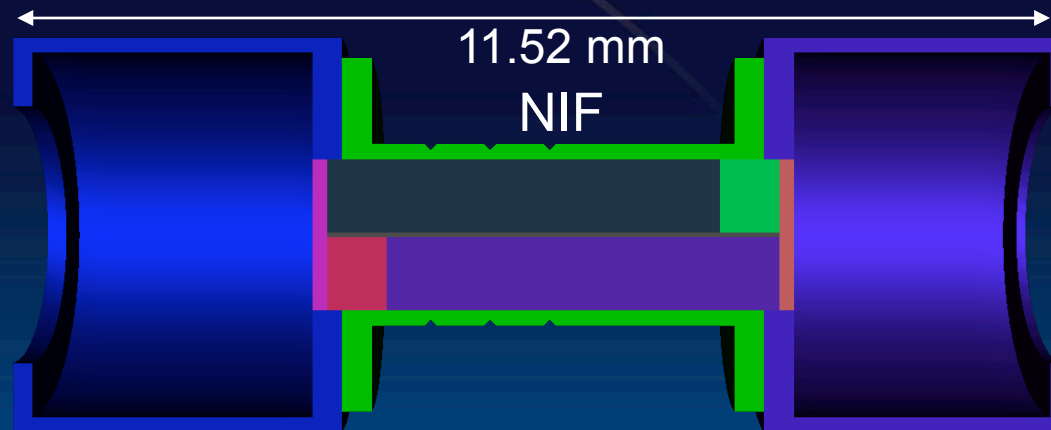
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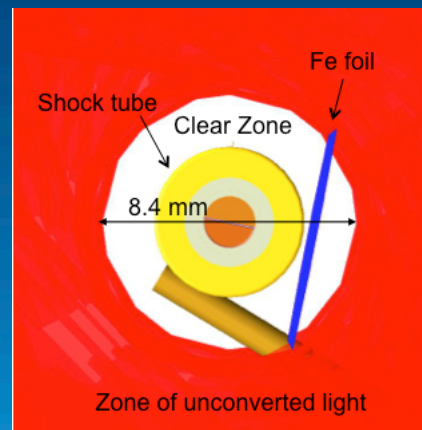


Omega

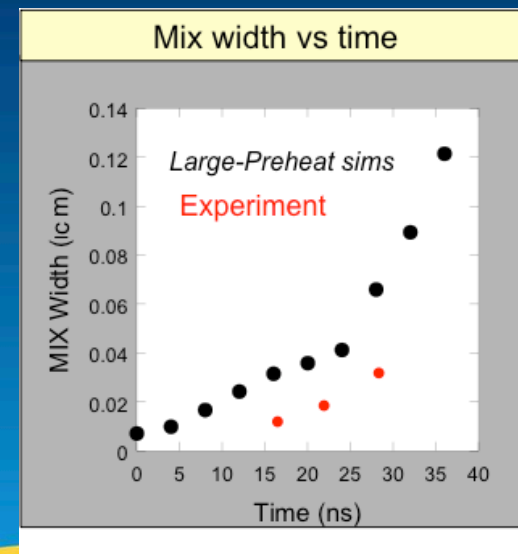


Developed BABL, which will be of general interest and utility:

- 3mm by 2mm spot
- 6 quads
- 3% conversion efficiency



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Opportunities

- Why does symmetry seem to matter more in experiments than in simulations of the low foot target? Is this ameliorated by increasing the adiabat?
- The velocities we are seeing at 'stagnation' make me question our picture about what is going on in the end game.
- Chunks of material that find themselves in the hot fuel evolve into structures small enough such that T in the fuel can find D in the CD ablator. How does that process go forward? Is BHR 3 the right answer?

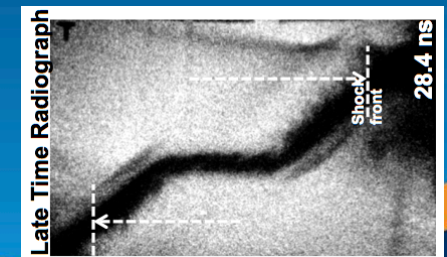
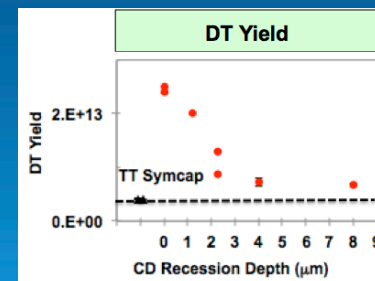
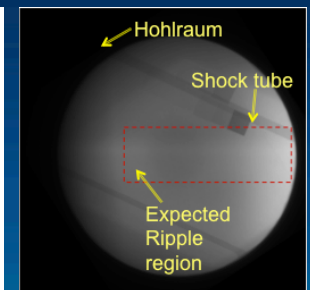
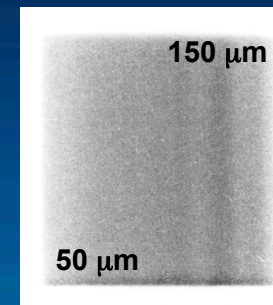
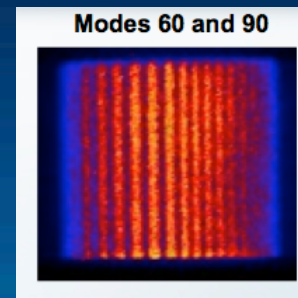
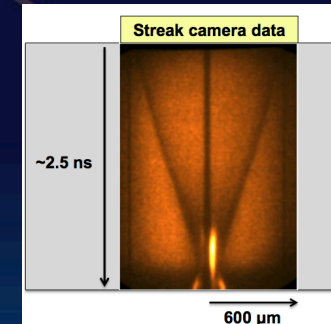
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